

# UV space telescope for astrophysical tests of redshift drift and the stability of fine-structure constant

---

CATARINA ALVES, TOMÁS SILVA

# Introduction

---

- Astrophysical tests of the stability of fundamental constants

# Introduction

---

- Astrophysical tests of the stability of fundamental constants
  - high-resolution optical/UV spectroscopic measurements

# Introduction

---

- Astrophysical tests of the stability of fundamental constants
  - high-resolution optical/UV spectroscopic measurements
- Equivalence Principle violations

# Introduction

---

- Astrophysical tests of the stability of fundamental constants
  - high-resolution optical/UV spectroscopic measurements
- Equivalence Principle violations
  - Varying fundamental constants induce violations of the universality of free fall

# Introduction

---

- Astrophysical tests of the stability of fundamental constants
  - high-resolution optical/UV spectroscopic measurements
- Equivalence Principle violations
  - Varying fundamental constants induce violations of the universality of free fall
  - How to look for a varying constant ?
    1. One can only measure dimensionless combinations of dimensional constants
    2. Such measurements are necessarily local

# Introduction

---

- Astrophysical tests of the stability of fundamental constants
  - high-resolution optical/UV spectroscopic measurements
- Equivalence Principle violations
  - Varying fundamental constants induce violations of the universality of free fall
  - How to look for a varying constant ?
    1. One can only measure dimensionless combinations of dimensional constants
    2. Such measurements are necessarily local

$$\alpha = \frac{1}{4\pi\epsilon_0} \frac{e^2}{\hbar c}$$

# Introduction

---

- Astrophysical tests of the stability of fundamental constants
  - high-resolution optical/UV spectroscopic measurements
- Equivalence Principle violations
  - Varying fundamental constants induce violations of the universality of free fall
  - How to look for a varying constant ?
    1. One can only measure dimensionless combinations of dimensional constants
    2. Such measurements are necessarily local

$$\boxed{\alpha} = \frac{1}{4\pi\epsilon_0} \frac{e^2}{\hbar c}$$



# Introduction

---

- Astrophysical tests of the stability of fundamental constants
  - high-resolution optical/UV spectroscopic measurements
- Equivalence Principle violations
  - Varying fundamental constants induce violations of the universality of free fall
  - How to look for a varying constant ?

## **Current data**

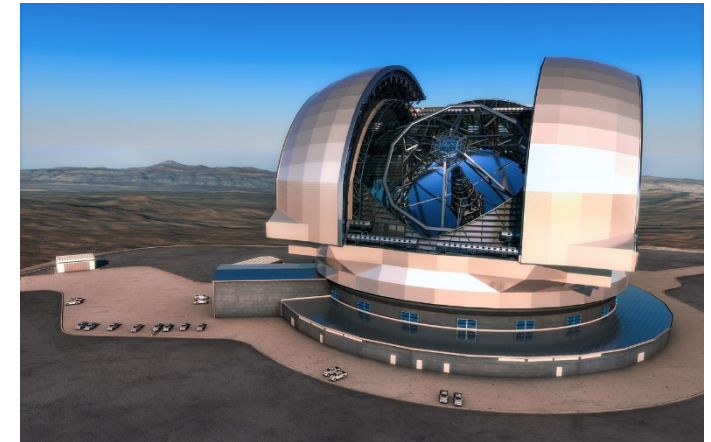
- Useful constraints

# Introduction

---

- Astrophysical tests of the stability of fundamental constants
  - high-resolution optical/UV spectroscopic measurements
- Equivalence Principle violations
  - Varying fundamental constants induce violations of the universality of free fall
  - How to look for a varying constant ?

**Current data → ELT (HIRES)**



Credit: ESO

Artist's impression of the European Extremely Large Telescope

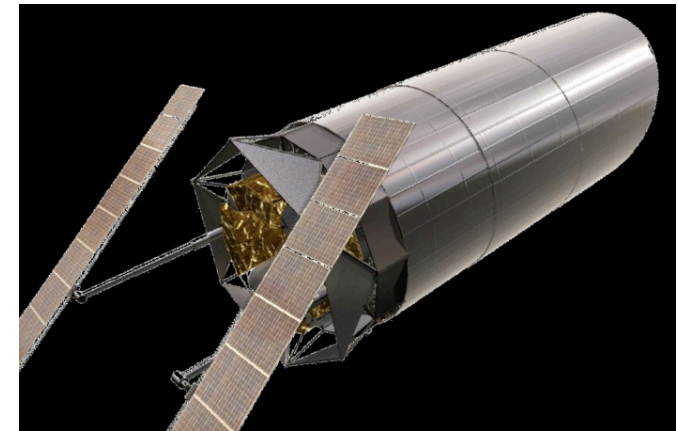
# Introduction

---

- Astrophysical tests of the stability of fundamental constants
  - high-resolution optical/UV spectroscopic measurements
- Equivalence Principle violations
  - Varying fundamental constants induce violations of the universality of free fall
  - How to look for a varying constant ?

**Current data → ELT (HIRES) → Space**

Credit: MSFC Advanced Concepts Office



8m ATLAST artist's conception  
(STSCI)

# Redshift

---

•

Credit: Dr. Dan Russell, Grad. Prog. Acoustics, Penn State.

# Redshift

---

•

Credit: Dr. Dan Russell, Grad. Prog. Acoustics, Penn State.

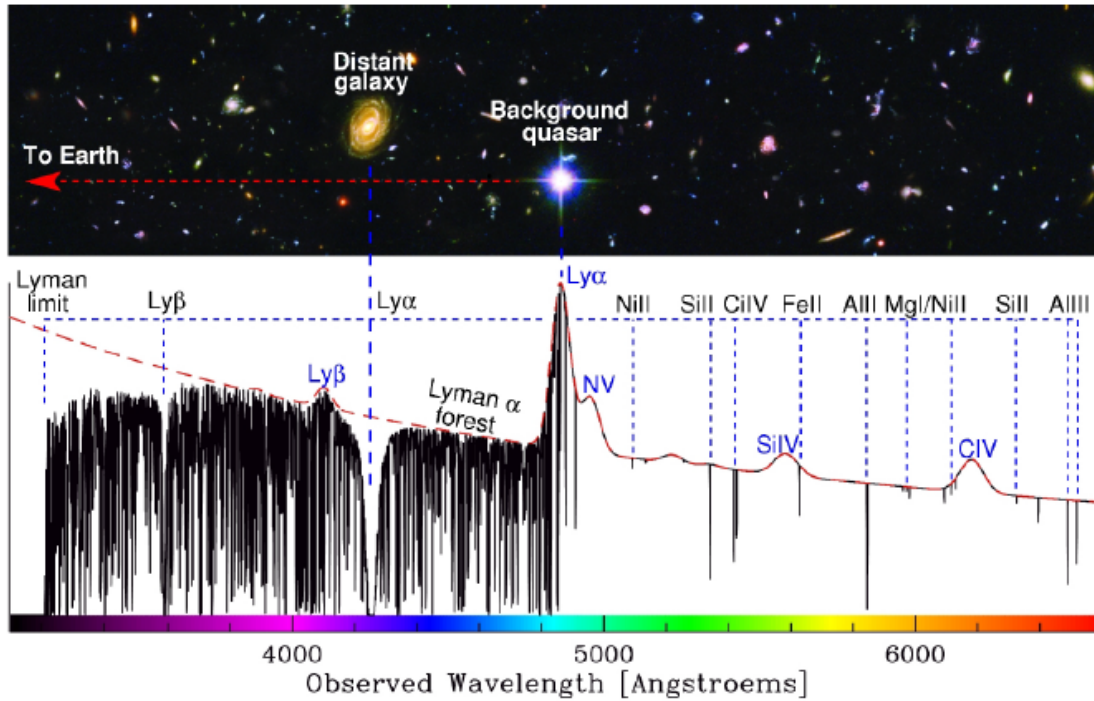
# Redshift

---

$$z = \frac{\lambda_{obs} - \lambda_{em}}{\lambda_{em}}$$

Credit: Dr. Dan Russell, Grad. Prog. Acoustics, Penn State.

# Redshift

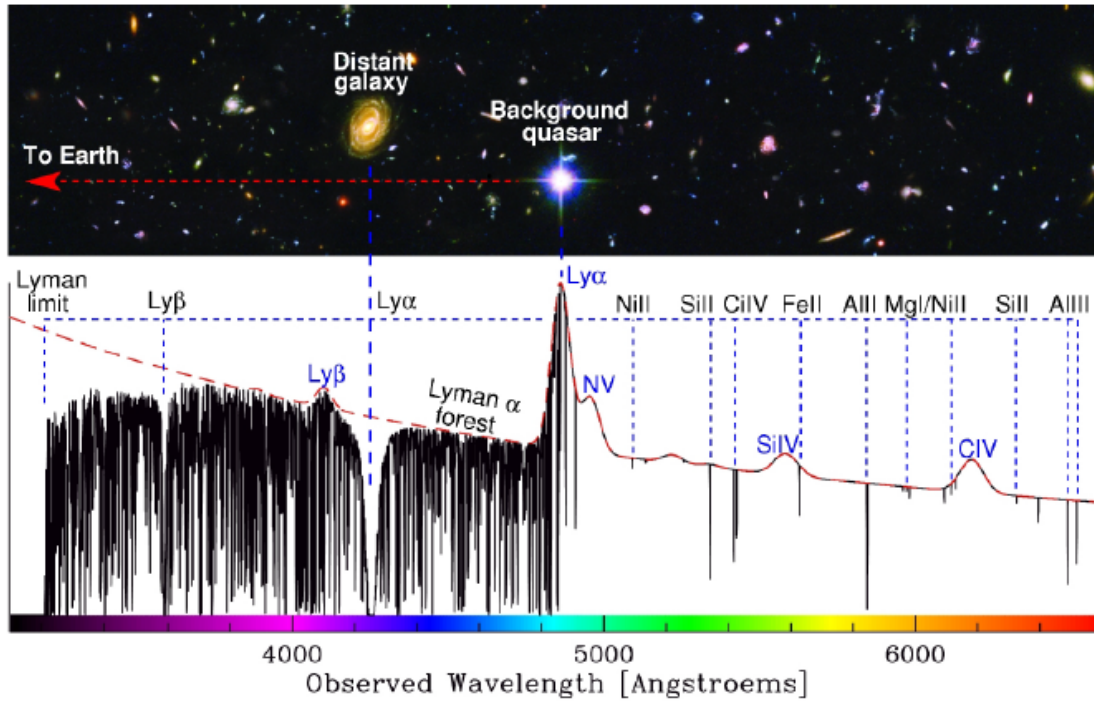


Credit: Michael Murphy, Swinburne University of Technology, Melbourne, Australia.

$$z = \frac{\lambda_{obs} - \lambda_{em}}{\lambda_{em}}$$

$$\frac{dz}{dt_0} = \frac{d}{dt_0} \left[ \frac{a(t_0)}{a(t_e)} - 1 \right] = \frac{\dot{a}(t_0) - \dot{a}(t_e)}{a(t_e)} \approx (1+z)H_0 - H(z)$$

# Redshift



Credit: Michael Murphy, Swinburne University of Technology, Melbourne, Australia.

$$z = \frac{\lambda_{obs} - \lambda_{em}}{\lambda_{em}}$$

$$\frac{\Delta\alpha}{\alpha} = \frac{\alpha(z) - \alpha_0}{\alpha_0}$$



# Varying $\alpha$

---

- Assumptions made to relate the evolution of  $\alpha$  to that of dark energy.

$$\frac{\Delta\alpha}{\alpha} = \zeta\kappa(\phi - \phi_0) \quad \Omega_\phi(z) \equiv \frac{\rho_\phi(z)}{\rho_{\text{tot}}(z)} \simeq \frac{\rho_\phi(z)}{\rho_\phi(z) + \rho_m(z)} \quad 1 + w_\phi = \frac{(\kappa\phi')^2}{3\Omega_\phi}$$

$$\frac{\Delta\alpha}{\alpha}(z) = \zeta \int_0^z \sqrt{3\Omega_\phi(z') [1 + w_\phi(z')]} \frac{dz'}{1 + z'}$$

# Fiducial models

---

- Constant equation of state

$$w \neq 0 = \text{const}$$

# Fiducial models

---

- Constant equation of state

$$w_0 = \text{const}$$

- Dilaton-type model

$$w(z) = \frac{[1 - \Omega_\phi (1 + w_0)] w_0}{\Omega_m (1 + w_0) (1 + z)^3 [1 - \Omega_\phi (1 + w_0)] - w_0}$$

$$\phi(z) \propto (1 + z) \quad \Omega_\phi(z) [1 + w(z)] = \text{const.}$$

# Fiducial models

---

- Constant equation of state

$$w_0 = \text{const}$$

- Dilaton-type model

$$w(z) = \frac{[1 - \Omega_\phi (1 + w_0)] w_0}{\Omega_m (1 + w_0) (1 + z)^3 [1 - \Omega_\phi (1 + w_0)] - w_0}$$

$$\phi(z) \propto (1 + z) \quad \frac{\Delta\alpha}{\alpha}(z) = \zeta \sqrt{3\Omega_\phi(1 + w_0)} \ln(1 + z)$$

# Fiducial models

---

- Constant equation of state

$$w_0 = \text{const}$$

- Dilaton-type model

$$w(z) = \frac{[1 - \Omega_m (1 + w_0)] w_0}{\Omega_m (1 + w_0) (1 + z)^3 [1 - \Omega_m (1 + w_0)] - w_0}$$

- Chevallier-Polarski-Linder (CPL) parametrization

$$w(z) = w_0 + w_a z / (1 + z)$$

# Parameters

---

- $\Omega \downarrow m, fid = 0,3$   $\sigma \downarrow \Omega \downarrow m = 0,03$
- $w \downarrow 0, fid = -0,9$   $\sigma \downarrow w \downarrow 0 = 0,1$
- $w \downarrow a, fid = 0,3$   $\sigma \downarrow w \downarrow a = 0,3$
- $\zeta \downarrow fid = 0 ; 5 \cdot 10 \uparrow -6$   $\sigma \downarrow \zeta = 10 \uparrow -4$

# Parameters

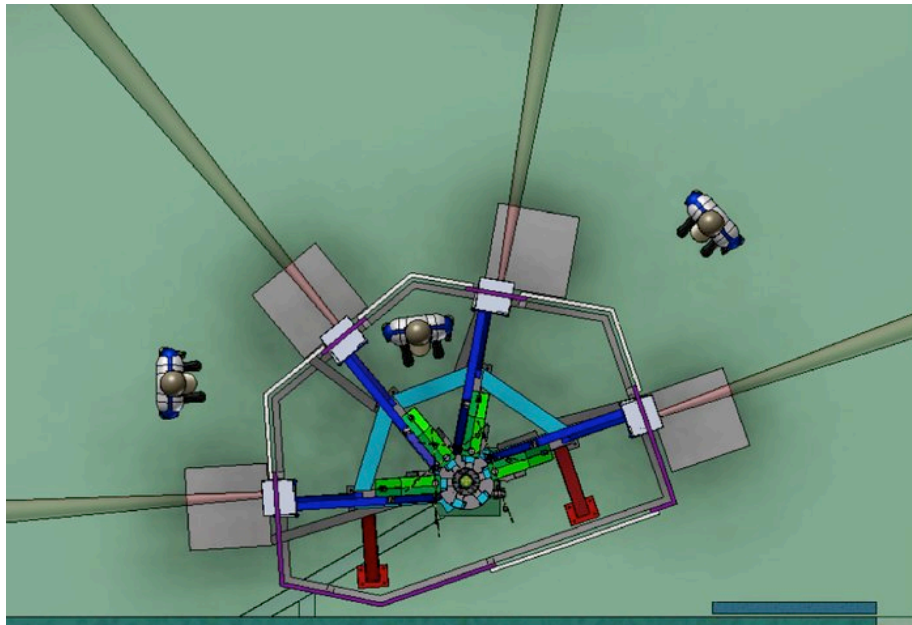
---

- $\Omega_{m, fid} = 0,3$   $\sigma_{\Omega_{m, fid}} = 0,03$
- $w_{0, fid} = -0,9$   $\sigma_{w_{0, fid}} = 0,1$
- $w_a, fid = 0,3$   $\sigma_{w_a, fid} = 0,3$
- $\zeta_{fid} = 0 ; 5 \cdot 10^{-6}$   $\sigma_{\zeta} = 10^{-4}$

Cosmological datasets

# Important remarks

---



Combining the light of four VLTs for the ESPRESSO instrument  
Credit: ESO/ESPRESSO Consortium

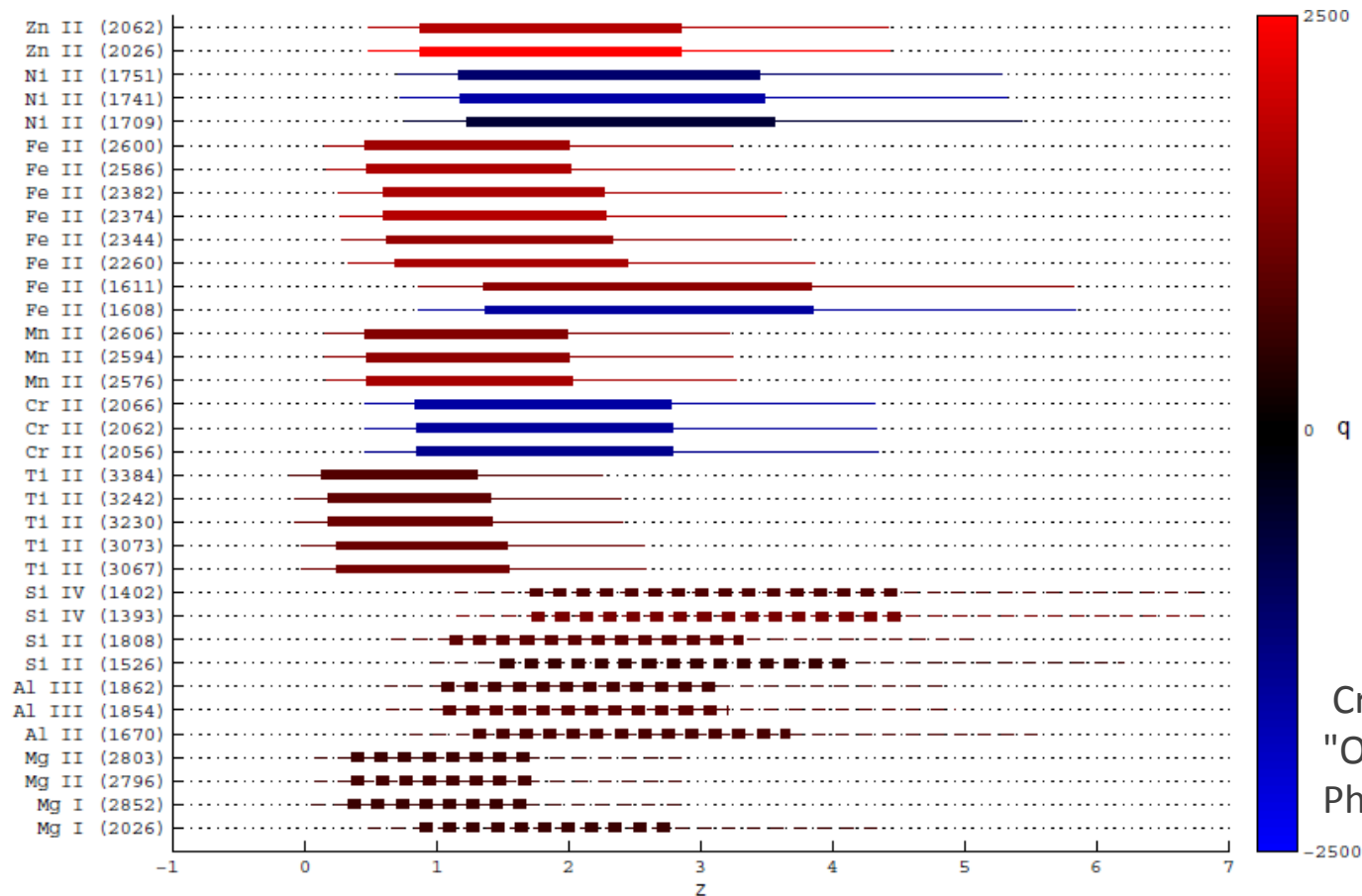
- Improvement of current bounds on the Eötvös parameter

ESPRESSO vs MICROSCOPE

HIRES vs STEP



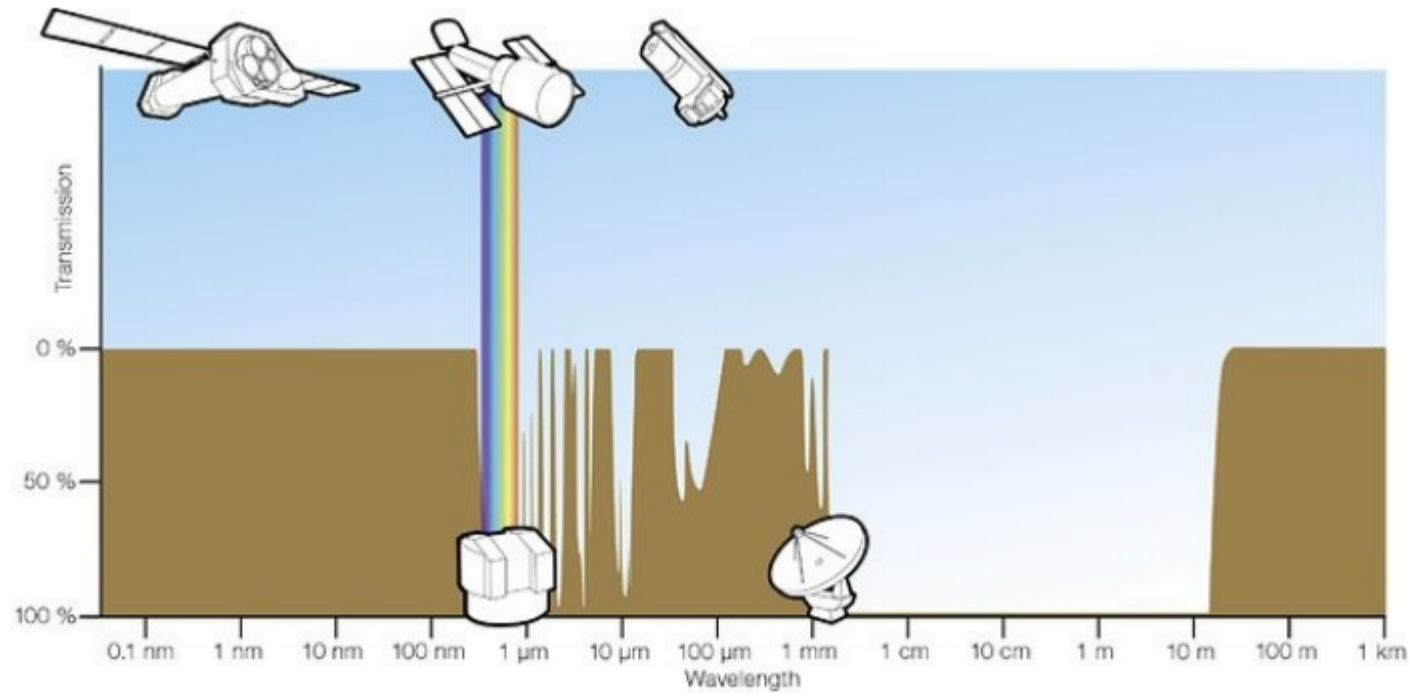
# Measurements of the fine-structure constant



$$\Delta v \approx -\frac{2cq_i}{\omega_0} \left( \frac{\Delta \alpha}{\alpha} \right)$$

Credit: de Oliveira Leite, Ana Catarina.  
"Optimization of ESPRESSO Fundamental  
Physics Tests." (2015)

# UV advantages



Credit: ESA / Hubble / F Granato

# Analysis

- Targets (quasars)
- Recuperation of fiducial models
- Uncertainty Earth vs space

```
def Fisher(zS, sigmas, Hf, h, Om, w0, wa, N=1): # matrizes de Fisher de tama
    '''sigmas e zS têm o mm tamanho; Hf-funcao a usar; h=0, Om=1, w0=
    n=4 # está adaptada para isto
    F = np.zeros([n,n])
    for b in range(len(zS)):
        elem = np.ones([n,n]) # inicializar a matriz a somar
        for i in range(n):
            if i==0:
                deriv = dfdh(Hf, zS[b], h, Om, w0, wa, N)
            elif i==1:
                deriv = dfdOm(Hf, zS[b], h, Om, w0, wa, N)
            elif i==2:
                deriv = dfdw0(Hf, zS[b], h, Om, w0, wa, N)
            else:
                deriv = dfdwa(Hf, zS[b], h, Om, w0, wa, N)
            elem[:,i]*=deriv # preencher a linha i
            elem[i,:]*=deriv # preencher a coluna i
        elem*=sigmas[b]**-2
        F=F+elem
    if all(F[:,3]==0)==True and all(F[3]==0)==True: # não tem wa
        F=np.delete(F, 3, 1) # tira 4ª coluna
        F=np.delete(F, 3, 0) # tira 4ª linha
    if all(F[:,2]==0)==True and all(F[2]==0)==True: # não tem w0
        F=np.delete(F, 2, 1)
        F=np.delete(F, 2, 0)
    return F
#Fim Construir F#
```

# Targets

---

- Data set – 333 quasars
- 35 transition
- Select the right targets:
  - wavelength
  - $\alpha$  variations: anchor, redshifter, blueshifter

# Recuperate the models

---

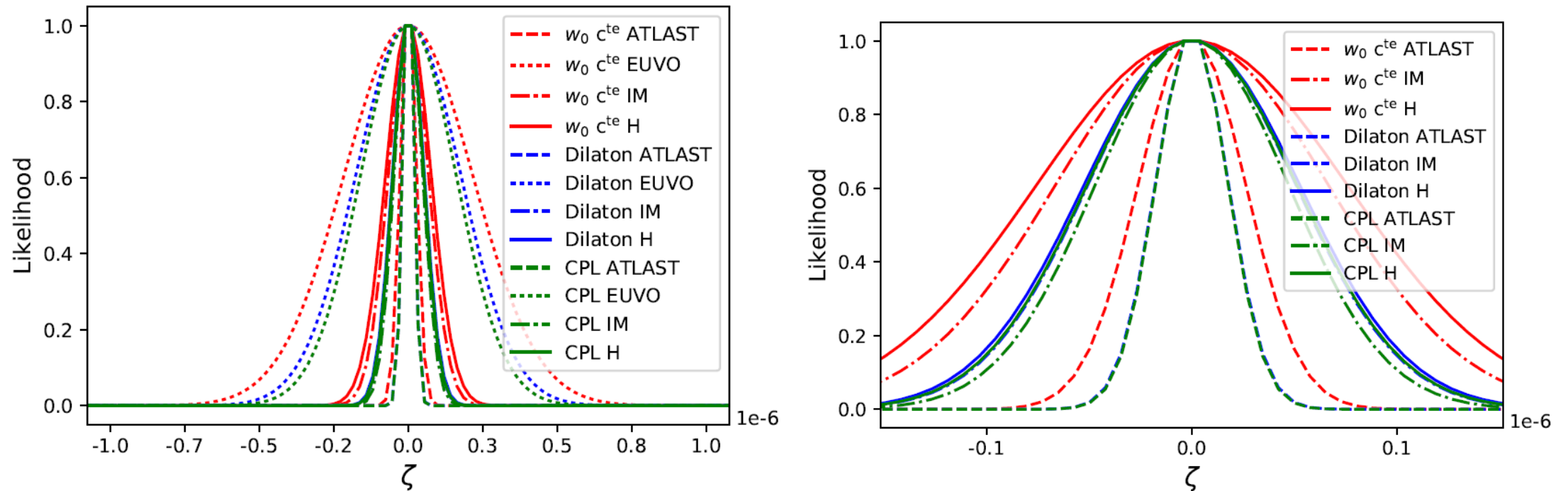
- 3 study cases:
  - ATLAST:  $\lambda = [1100 - 25000] A$
  - EUVO:  $\lambda = [2400 - 3500] A$
  - Imaginary case:  $\lambda = [3500 - 4000] A$
- Uncertainties → Earth telescope ELT - HIRES
- Fisher matrix techniques

# Fisher Matrix

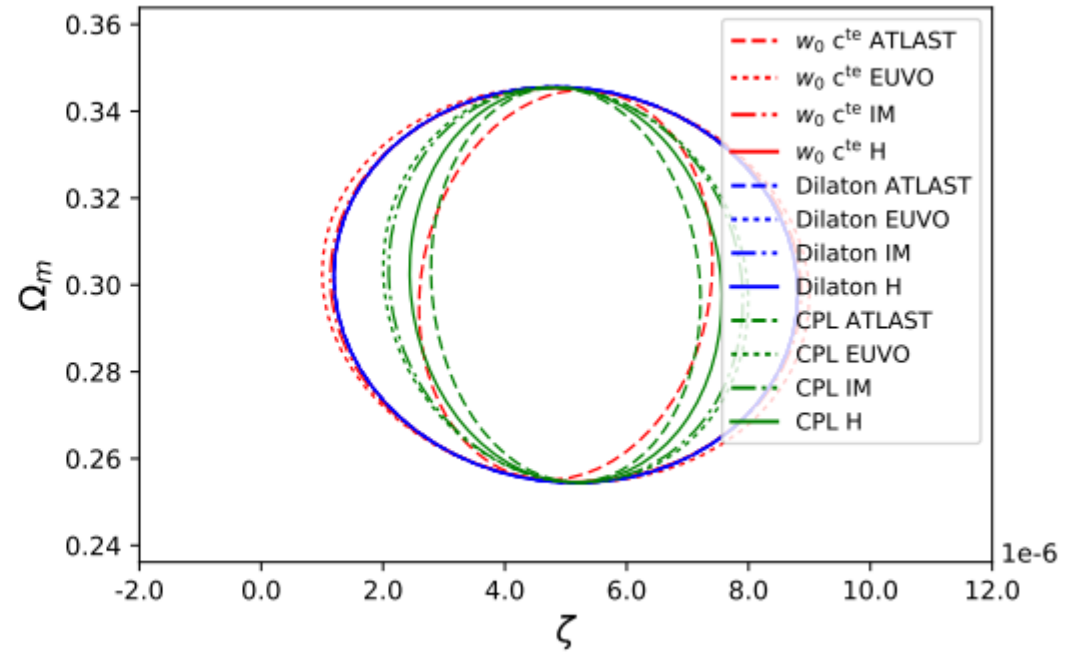
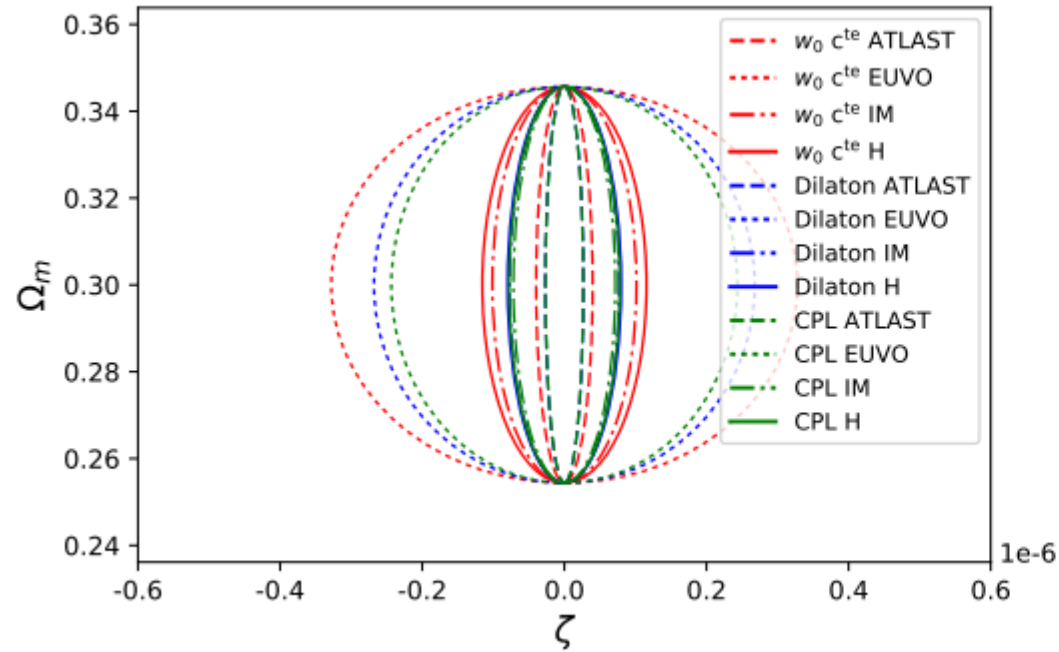
---

- Forecast the precision and impact of future experiments
- Analyze the combination of cosmological constraints from various data sets
- Experimental design
- Fisher matrix is the inverse of the covariance matrix
$$Cov(X,Y)=\langle (X-x)(Y-y) \rangle$$

# Results – Fine-structure

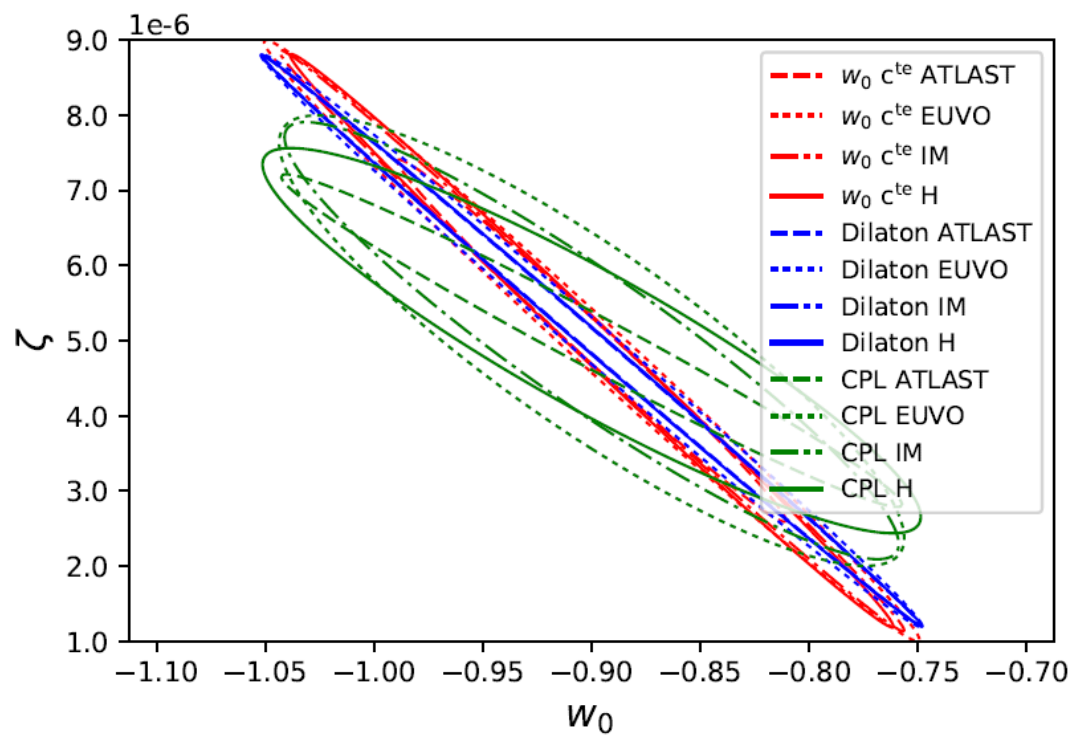
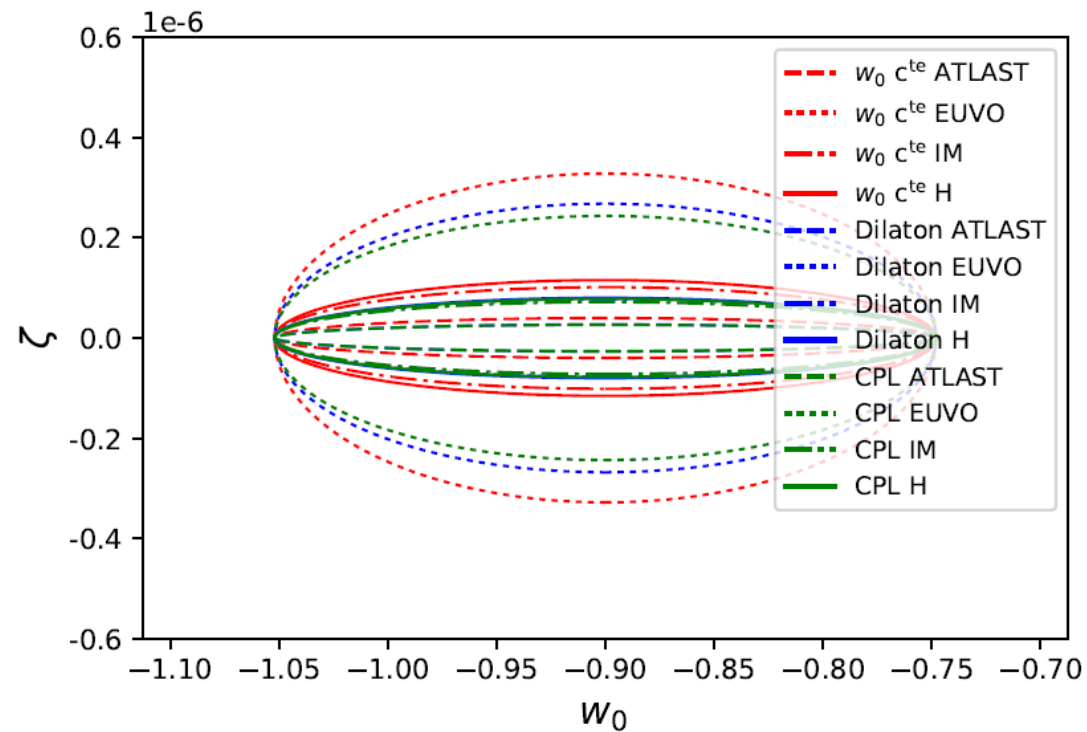


# Results – Fine-structure

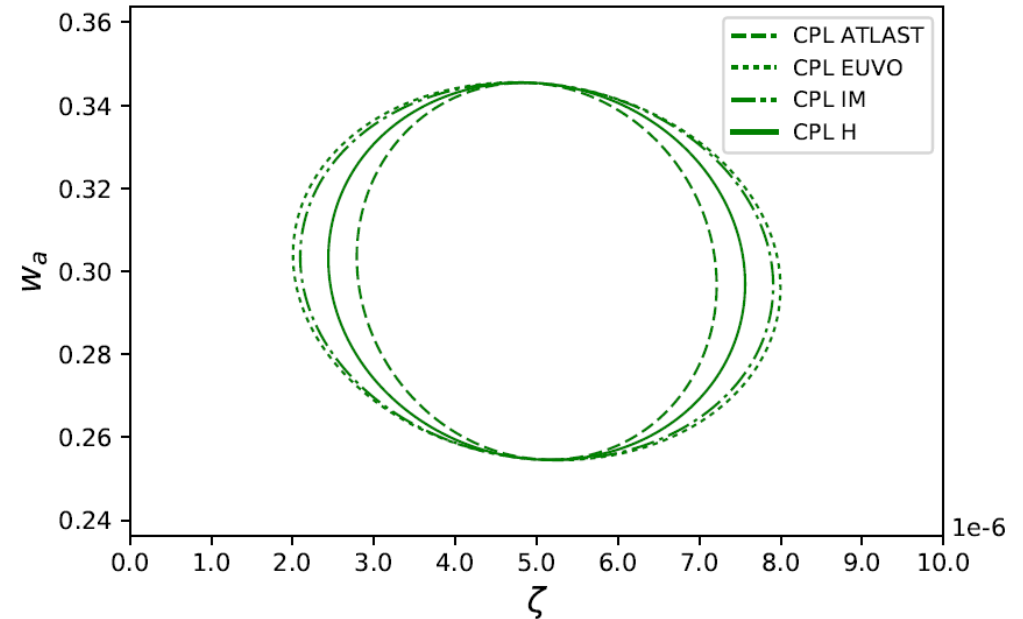
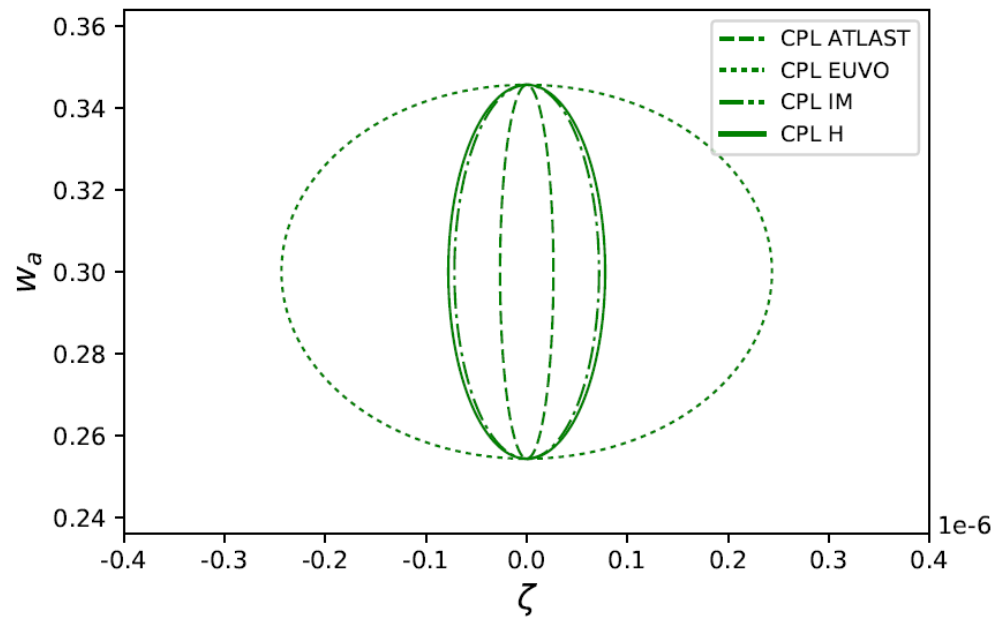




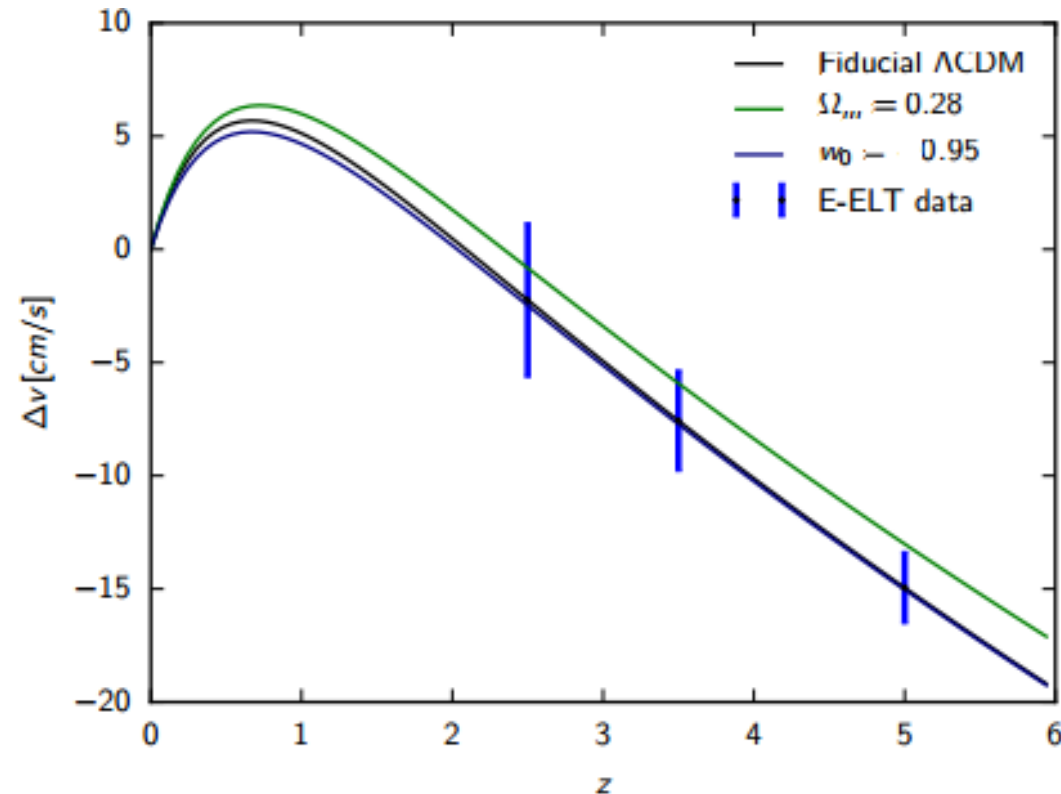
# Results – Fine-structure



# Results – Fine-structure

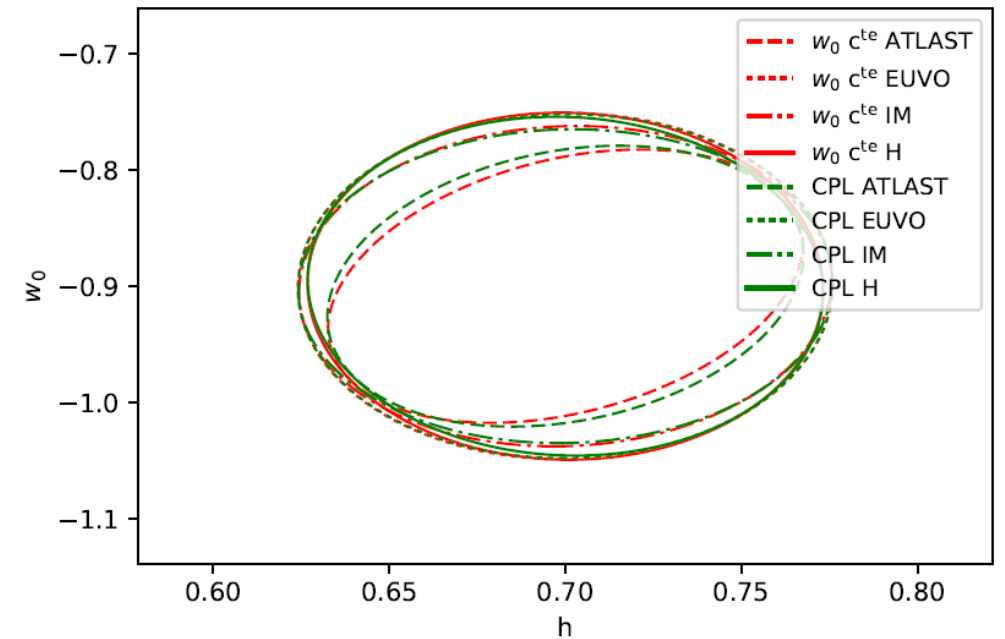
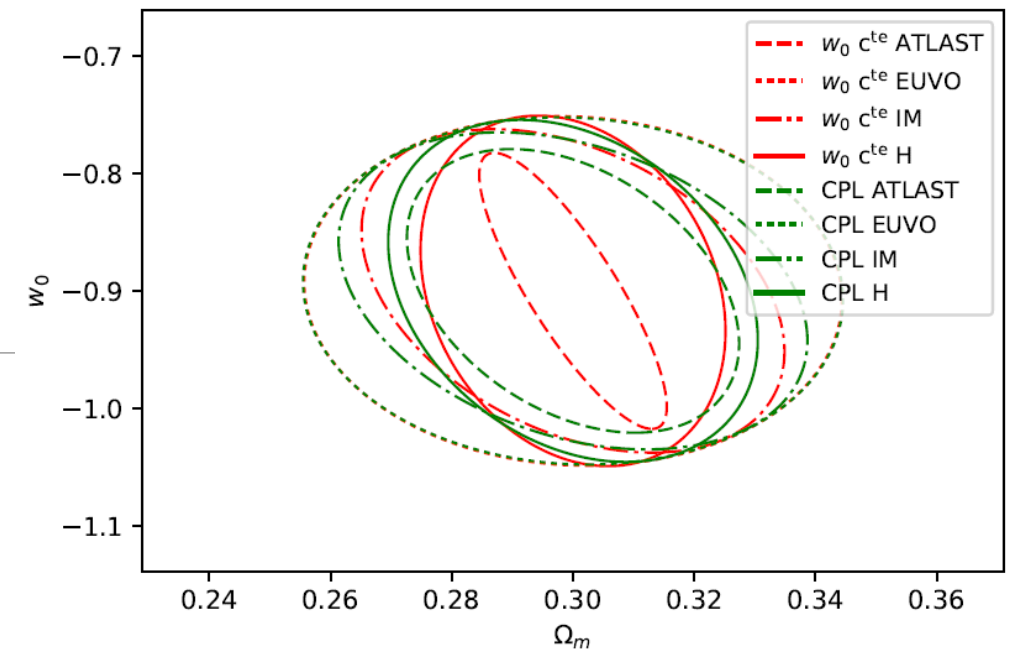
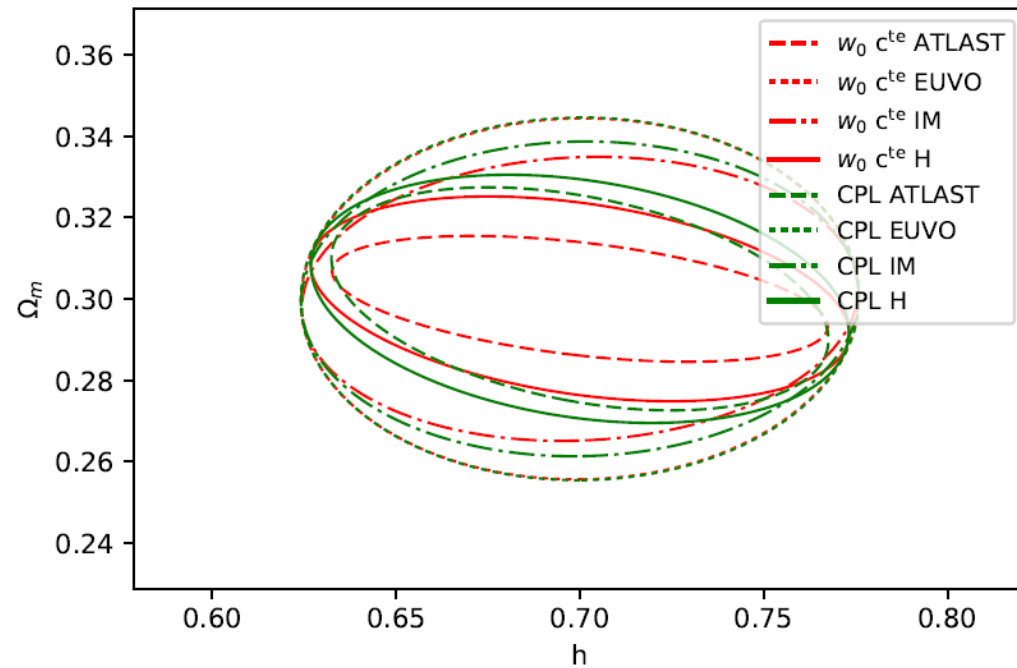


# Redshift drift



Credit: Martins et al. 2016

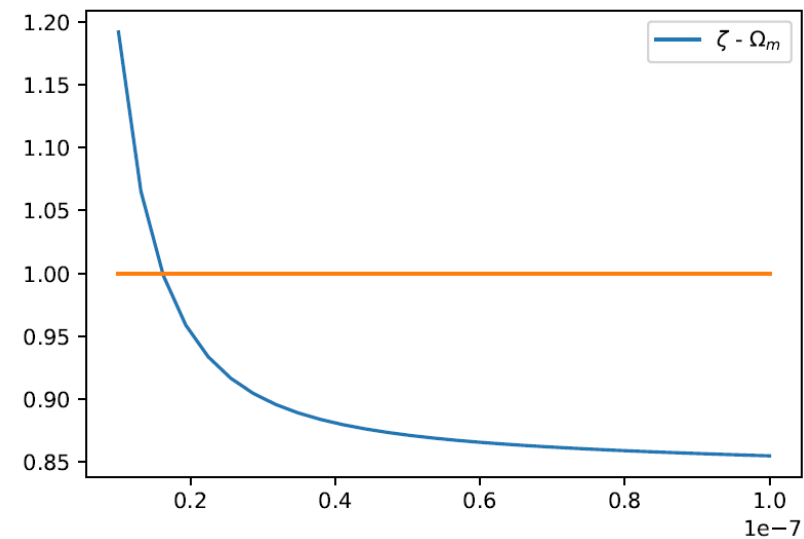
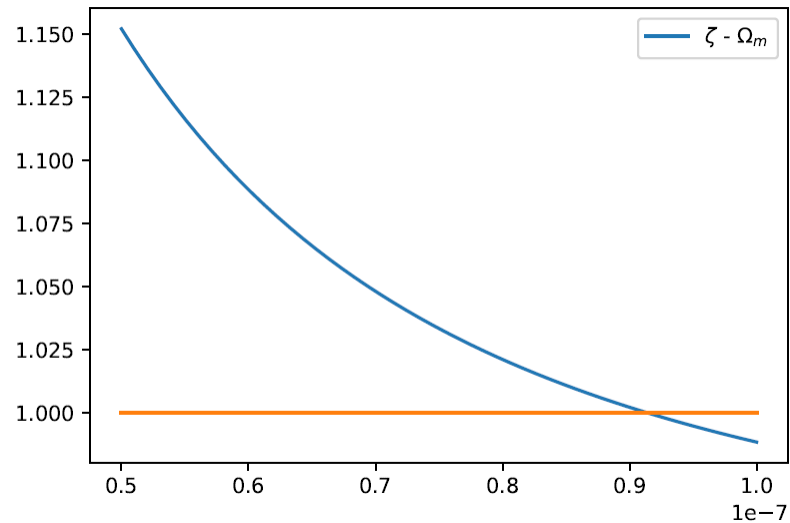
# Results – Redshift drift



# Uncertainty Earth vs space

- Figure of merit

$$FOM = 1/area$$



# Conclusions

---

- UV is a very important frequency and it is useful to observe
- ATLAST is the best proposal now
- ATLAST provide better constrains to cosmological parameters
- Suggestion: modifications to EUVO proposal
- Suggestion: space spectroscope with  $\lambda=[3500-4000] \text{ \AA}$
- Better constrains: join measurements for  $\Delta\alpha/\alpha$  and redshift drift

# Future

---

- Different experiments to get different correlations → break degeneracies
- Our results would have been better if there were more available transitions → this is an area that should be worked on → has potential for lots of scientific knowledge on it